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The SeaWinds scatterometer, launched onboard the QuikSCAT satellite in 1999, measures global ocean vector winds. In addition to measuring radar backscatter, SeaWinds simultaneously measures the microwave brightness temperature of the atmosphere/surface, and this passive microwave measurement capability is known as the QuikSCAT Radiometer (QRad). This paper presents a QRad retrieval algorithm used to infer instantaneous oceanic rain rates. This statistical algorithm is trained using near-simultaneous observations of major rain events by QRad and the Tropical Rainfall Measuring Mission (TRMM) Microwave Imager (TMI). Rain rate retrieval algorithm validation is presented through comparisons with independent rain measurements from the TMI 2A12 surface rain rates and the TRMM 3B42RT composite microwave and visible and infrared near-real time data product. Results demonstrate that QRad rain rate measurements are in good agreement with these independent microwave rain observations and superior to the visible/infrared rain estimates. Thus the QRad rain measurement time series is a valuable addition to the oceanic precipitation climatology that can be used to improve the diurnal estimation of the global rainfall, which is a goal for the future Global Precipitation Mission program. Moreover, the availability of QRad data will provide GPM users early access to learn to use less-precise rain measurements that will occur in the GPM era with the use of less-capable constellation satellites. Finally, these QRad rain estimates will be available in the planned data reprocessing (FY 2006) of QuikSCAT winds to improve the rain flagging of rain-contaminated oceanic wind vector retrievals.


1. Introduction

[2] For more than one decade, multi-frequency microwave radiometer imagers flying on low earth satellites have provided valuable day/night remote sensing of oceanic and atmospheric variables; but the emphasis on oceanic precipitation measurements achieved a significant advance with the launch of the Tropical Rainfall Measuring Mission (TRMM) observatory in late 1997. Because of TRMM’s non-sun synchronous orbit, for the first time, precipitation measurements were available from a satellite over all local times so that the diurnal cycle of oceanic precipitation could be studied. However, from 1998 through late 2002, the ocean sampling was very sparse with only four such satellite instruments operating on-orbit; three Defense Meteorological Support Program (DMSP) satellites carrying the Special Sensor Microwave Imager (SSM/I), and the Tropical Rainfall Measuring Mission’s (TRMM) Microwave Imager (TMI). The SSM/I’s fly on near-polar sun synchronous satellites that provide greater than 90% earth coverage daily; however, since they fly in a day/night terminator orbit, they provide only morning and evening sampling times. On the other hand, the TMI flies in a low inclination (38°) non-sun synchronous orbit that has been optimized to measure tropical rainfall. TMI provides full diurnal sampling over the period of slightly greater than one month. However, even with the four passive microwave sensors, the statistics of oceanic rainfall were badly under-sampled. Since the fall of 2002, a fifth microwave imager, the Advanced Microwave Scanning Radiometer (AMSR-E) on NASA’s Aqua earth observing system satellite began its ocean precipitation measurements; but even with this additional radiometer, the diurnal sampling is still less than desired.

[3] Many researchers [e.g., Wilheit et al., 1991; Petty and Katsaros, 1992; Bell and Reid, 1993; Chang et al., 1995; Imaoka and Spencer, 2000] have studied diurnal sampling...
of oceanic precipitation using satellite microwave radiometers. Because of the sparse sampling, diurnal cycles must be estimated using large space-time averages, and likewise, it is difficult to determine the rainfall statistics for regional oceanic precipitation. In the future, a constellation of satellites, known as the Global Precipitation Mission (GPM) \cite{Smith_2001, Smith_2001}, will solve this observational shortage. This proposed constellation, comprised of satellites in low inclination and polar low-earth orbits (non-sun synchronous and sun synchronous) will provide near-global coverage with a worst case revisit time of three hours at the equator. An important aspect of GPM is the use of a highly capable “core observatory” (similar to TRMM) to provide rainfall classification and rain rate retrievals. This will be augmented by six or more less-capable “constellation” satellites carrying microwave radiometers, which are cross-calibrated to the core observatory, and provide the rapid temporal sampling of rainfall. Thus, in the future, scientists and operational users will have to learn to accommodate rain retrievals of varying quality in their research and applications.

In September 1999, the QuikSCAT Radiometer (QRad) began ocean precipitation measurements, which provides additional independent samples over SSM/I and TMI. A typical example of the QRad sampling is shown in Figure 1 for a three-hour window (universal time: 00:00–03:00). Also shown are the corresponding sampling coverage for TMI and three SSM/I’s. It is observed that QRad increases the coverage area by about 10%; but even with five microwave imagers, the ocean sampling is still only approximately 60% in a typical 3-hour window. Never the less, the QRad’s sampling contribution is significant in that the daily average revisit time is reduced as shown in Figure 2. For clarity of presentation, sampling improve-
ments, due to the QRad and averaged over 20° latitudinal zones, are quantitatively summarized in Table 1. Further, an additional illustration of QRad oceanic sampling contribution is shown in Figure 3 which presents a typical scenario of "local time of day" QRad sampling over a 1° × 1° box located at equator and prime meridian for a period of one month. Also shown are the local time samplings for TMI and three SSM/I instruments. It is clear that QRad is providing independent sampling which complements and fills in the gaps between the sampling times of the other satellites.

Thus the QRad time series (from September 1999 to present) is a valuable addition to the ocean precipitation climate data set. Further, the early availability of QRad rain measurements provides an excellent opportunity for learning how to utilize future GPM data sets. As will be described, even though the quality of the QRad rain retrievals are somewhat limited compared to TMI and SSM/I, they certainly are useful in that they provide additional temporal/spatial sampling. Moreover, they provide simultaneous, collocated precipitation measurements with QuikSCAT ocean surface wind vectors for rain-flagging contaminated wind vector retrievals.

![Figure 3. Typical time of day sampling for SSMI (F-13, F-14 and F-15), TMI and QuikSCAT Radiometer. Sample location is 1° × 1° latitude/longitude box located at equator and prime meridian.](image1)

Table 1. Average Oceanic Coverage in a Typical 3-Hour Window

<table>
<thead>
<tr>
<th>Region</th>
<th>Ocean Coverage Without QRad</th>
<th>Ocean Coverage With QRad</th>
<th>QRad Contribution</th>
</tr>
</thead>
<tbody>
<tr>
<td>40°N–60°N</td>
<td>57.94 %</td>
<td>68.55 %</td>
<td>10.61 %</td>
</tr>
<tr>
<td>20°N–40°N</td>
<td>58.23 %</td>
<td>64.69 %</td>
<td>6.46 %</td>
</tr>
<tr>
<td>0°–20°N</td>
<td>63.70 %</td>
<td>71.34 %</td>
<td>7.64 %</td>
</tr>
<tr>
<td>20°S–0°</td>
<td>63.13 %</td>
<td>70.17 %</td>
<td>6.04 %</td>
</tr>
<tr>
<td>40°S–20°S</td>
<td>57.69 %</td>
<td>63.87 %</td>
<td>6.18 %</td>
</tr>
<tr>
<td>60°S–40°S</td>
<td>58.34 %</td>
<td>68.26 %</td>
<td>9.93 %</td>
</tr>
</tbody>
</table>

*Improvements due to QRad contribution are calculated for regions of 20° latitudinal zones.

![Figure 4. Measurement geometry of SeaWinds instrument onboard QuikSCAT satellite.](image2)
Figure 5. Brightness temperature spectral ratio as a function of columnar water vapor. (top) A plot of horizontal polarization and (bottom) the vertical polarization. Circles denote binned/averaged data and the error bars show ± one standard deviation. The solid line shows the third order polynomial fit.
Figure 6. Comparison of QRad and TMI ocean brightness temperatures for rain-free five day averages. Circles are binned/averaged data, and error bars represent ± one standard deviation. Dashed line is perfect agreement and solid line shows least squares regression.

Figure 7. Five-day average oceanic brightness temperature differences (QRad – TMI) for rain-free ocean, April 2003. Circles are binned/averaged in 5 K bins by TMI, and error bars denote ± one standard deviation.
In this paper, the oceanic rain measurements made with QRad are described. Section 2 describes the QuikSCAT instrument and the external radiometric calibration procedure. The QRad rain rate algorithm is discussed in detail in section 3; and validation of QRad rain measurements through comparisons with other independent rain measuring instruments are presented in section 4. Results demonstrate that the QRad oceanic rain estimates are in good agreement with TMI and SSMI independent rain measurements.

2. QuikSCAT Radiometer

2.1. Instrument Description

The SeaWinds scatterometer on the QuikSCAT satellite is a conical scanning long-pulse radar system used to measure the backscatter from the ocean surface to infer surface wind speed and direction [Spencer et al., 1997]. This scatterometer has two receiver channels, which allow the received backscatter signal (echo) and the black-body microwave emission (noise) from the ocean surface and interviewing atmosphere to be separated. Although, quantitative microwave brightness temperature measurements were not originally envisioned; never the less, the QuikSCAT radiometric function has been implemented post-launch through ground signal processing. Thus QRad measures the linearly polarized microwave brightness temperature $T_b$, at 13.4 GHz using a mechanical spinning reflector antenna as shown in Figure 4. Microwave emissions are collected over the entire conical scan (forward and aft looking) with separate offset “pencil beams” at 46° incidence (horizontal polarization, H-pol) and 54° incidence (vertical polarization, V-pol). Individual $T_b$'s are averaged on a spacecraft measurement grid of wind vector cells at 25 km resolution that results in mean horizontal and vertical $T_b$'s collocated with the normalized backscatter measurements. The pulse repetition frequency and antenna scan rate have been designed to provide approximately 50% overlap of the instantaneous field of view (IFOV) in both the along track and cross track directions. Thus with the two pencil beams, it is possible to isolate the microwave emissions from the earth into elliptical footprints defined by the one-way antenna pattern half-power contours (approximately 35 km x 50 km). Details of the QRad instrument and its radiometric calibration are provided by Jones et al. [2000] and Meershahi [2000].

2.2. Radiometric Calibration

Designed as a radar, SeaWinds is not an optimum radiometer. Brightness temperatures ($T_b$'s) are calculated for each received pulse with an equivalent integration time of 1.5 ms and a noise bandwidth of only 750 KHz. Because of the limited time-bandwidth product, the radiometric precision is much lower than desired ($\Delta T = 27$ Kelvin/pulse). For QRad rain measurements, this can be partially ameliorated by using spatial and temporal averaging where both for-

<table>
<thead>
<tr>
<th>Date</th>
<th>Offset</th>
<th>Slope</th>
</tr>
</thead>
<tbody>
<tr>
<td>Sept. 1999</td>
<td>6.55 K</td>
<td>0.977</td>
</tr>
<tr>
<td>June 2000</td>
<td>6.32 K</td>
<td>0.955</td>
</tr>
<tr>
<td>Jan. 2001</td>
<td>9.07 K</td>
<td>0.958</td>
</tr>
<tr>
<td>April 2003</td>
<td>4.67 K</td>
<td>0.978</td>
</tr>
</tbody>
</table>

aData are rain-free combined horizontal and vertical polarization three-day averaged ocean brightness temperatures. TMI brightness temperatures are interpolated to QRad frequency and extrapolated to QRad incidence angle.

Figure 8. Three-day average, rain-free, ocean brightness temperature probability density function, January 15–17, 2000.
ward-looking and aft-looking azimuth directions are collocated onto a 0.5° × 0.5° earth-located grid that is approximately equivalent to the QRad antenna surface resolution (50-km). Each polarized Tb observation is the average of about 24 pulses that results in a ΔT = 5 K.

Unfortunately for QRad there are no provisions for the usual two-point, hot and cold, absolute brightness temperature calibration. However, the QRad radiometric gain calibration is accomplished once per antenna scan using an internal ambient temperature (warm) load in the receiver; and the Tb offset is established one time, in an on-orbit calibration in 2000, using external comparisons with a well-known natural black-body sources (the Amazon rainforest) and with selected rain-free ocean Tb measurement comparisons with TMI.

For the ocean calibration, rain-free QRad polarized Tb’s are averaged for 3-days and are spatially collocated with TMI brightness measurements (over ±40° latitude on a 0.25° latitude × 0.25° longitude grid). Because the polarized ocean Tb’s change with frequency and because TMI does not have a 13.4 GHz channel, a translation of TMI brightness temperatures must be performed before direct comparisons are possible with QRad. For TMI, the two lowest frequency channels (10.7 and 19.4 GHz) bracket the QRad frequency at 13.4 GHz; however, the incidence angles do not match. The TMI incidence angle is 52.8° for all channels; whereas, for QRad, the inner (H-pol) beam is 46° and the outer (V-pol) beam is 54°. Thus, as described below, TMI Tb’s are interpolated over frequency and extrapolated over incidence angle to create QRad equivalent Tb’s, which are used to establish the QRad absolute radiometric offset.

Analysis has shown that this spectral ratio yields equivalent QRad Tb’s accurate to within a few Kelvin [Mehershahi, 2000; Jones et al., 2000]. However, for the given frequencies, this spectral ratio exhibits a nearly exponential dependence on atmospheric columnar water...
vapor as shown in Figure 5. To derive this spectral ratio, over 72,000 ocean $T_b$ points were simulated at each 10.9, 13.4 and 19.4 GHz using atmospheric and oceanic environmental parameters from SSMI F-13 and NOAA NCEP numerical weather analysis. The spectral ratio was then calculated at each $T_b$ location and binned and averaged in 2 mm water vapor bins represented by circles. The error bars denote ± one standard deviation. The natural logarithm of the spectral ratio was then regressed against water vapor using a third order polynomial fit shown by the solid line. Thus an estimate of the columnar water vapor, derived from collocated TMI retrievals, is used to select the proper value for the spectral ratio.

Further, because the orbital measurement swaths for QRad and TMI are not collocated simultaneously, transient rain events are present in both ocean data sets that can produce significant differences (10’s of Kelvin) at a given locations. This “error” is effectively removed by editing the data using TMI (and QRad) rain flags. If either instrument indicates rain, the location is deleted.

For land, the emissivity is more complex, and the radiation transfer model was not used to produce equivalent QRad $T_b$’s. However, the Amazon rain forest was used because it is a large isotropic and nearly homogeneous target that is an approximate blackbody with a brightness of about 285 K over this range of frequencies. Small diurnal effects of a few K have been observed in SSM/I measurements during ascending and descending pass times that are separated by approximately 12 hours, but during the 3-day average QRad Amazon comparisons, the TMI measured brightness temperatures at 10.7 and 19.4 GHz were averaged and linearly interpolated to compare with QRad $T_b$’s.

An example of the linear regression scatter diagrams for QRad and TMI equivalent $T_b$’s is given in Figure 6 for both H- and V-pols; and an expanded view of the difference between QRad and TMI measurements is shown in Figure 7. The symbols are binned average data on the TMI $T_b$; and the error bars denote ± one standard deviation. The stability of this external calibration procedure is good as observed from the resulting regression slope and offset for several different calibrations during 1999 to 2003 that are provided in Table 2.

Figure 10. Pacific Ocean brightness temperature deviation from the mean. Measurements are for repeating ground swaths, approximately four days separation.

Figure 11. QRad rain rate algorithm block diagram.
Another assessment of the calibration stability compares histograms of QRad and TMI equivalent ocean $T_b$'s taken seasonally. Here, three-day sets of average ocean brightness temperatures were produced with rain removed, and a typical set of histograms is shown in Figure 8. For H-pol, the QRad median $T_b$ is within a Kelvin of TMI; but for V-pol, the QRad results are low by a few Kelvin. Also QRad histograms are broader as the result of the increased QRad $\Delta T$. The year 2000 calibration statistics are tabulated in Table 3; and when taken over the year, the median differences show a slight systematic variation, which may be related to the QuikSCAT seasonal thermal environment. Over a period of one year, the global mean of this variation is $-0.29$ K with a standard deviation of 0.85 for horizontal and correspondingly $-2.76$ K with a standard deviation of 0.75 for vertical. Again these results demonstrate that QRad and TMI derived equivalent $T_b$ agree on average to within a few Kelvin.

The final example of relative $T_b$ stability is shown in the approximately two-year $T_b$ time series given in Figure 9. The object of this comparison is to assess whether or not there are variable $T_b$ biases caused by the seasonal solar heating of the satellite and instrument. This is important because the QRad transfer function uses the physical temperature of the front-end losses to calculate $T_b$. For this evaluation, the polarized brightness temperatures are averaged over all pixels for a repeating (every 4-day) ground swath in the middle of the Pacific ocean between ±45° latitude. During this evaluation, it was discovered that this orbit average $T_b$ is very stable even when rain pixels are included. Because both earth hemispheres (± latitudes) are included, the seasonal rain effects appear to cancel and the mean $T_b$ is very stable. In late 1999, a small step in $T_b$ is visible, which corresponds to a change in the QRad range gate width (equivalent to integration time); but since then there have been no changes in the instrument transfer function. In Figure 10, the QRad average polarized $T_b$ deviation from its polarized time series mean is displayed for these repeating ground tracks, and over this two-year period, the rms difference about the mean is 1.4 K for both polarizations. It is encouraging that both polarized brightness temperature deviations overlay and that they are consistent with the previous analysis presented above, which shows a small seasonal variation. These results demonstrate the stability and effectiveness of this external calibration technique used for QRad; and in fact, this external TMI calibration has produced very consistent results over the entire six years that QRad has been in operation.

Thus, based on analyses by Mehershahi [2000] and Jones et al. [2000], the absolute accuracy of the QRad $T_b$'s is estimated to be better than 4 K. While this does introduce a systematic bias error, its effects on rain retrieval are removed during algorithm training described below.

### 3. Integrated Rain Rate Algorithm

The QRad rain rate algorithm is a statistical based retrieval that uses an empirical brightness temperature - rain
rate ($T_b$-R) relationship to derive the integrated rain rate over the oceans [Ahmad et al., 2003]. Because the measured ocean brightness temperature is directly proportional to the path integrated rain rate, this is the chosen retrieved geophysical parameter. To calculate the average rain rate measured in mm/hr requires knowledge of the rain path length. Users may convert QRad integrated rain rate to surface rain rate by dividing by this rain path length that is equal to the height of the rain times secant (52.8°). The QRad $T_b$-R relationship was derived using a data set from rain events that were near-simultaneously observed by QRad and the TMI.

[20] A simplified algorithm block diagram is presented in Figure 11. The data inputs are (1) the QRad $T_b$’s from the QSCAT level 2A (L2A) and the retrieved wind speed from the QSCAT level 2B (L2B) data products available at http://podaac.jpl.nasa.gov/quickscat/ and (2) a priori information in the form of monthly-tabulated ocean background brightness temperatures.

[21] The individual polarized L2A QRad $T_b$’s and the L2B retrieved wind speed products are provided on a spacecraft measurement grid of wind vector cells at 25 km resolution. These two products are earth gridded and spatially averaged to 50 km resolution and used with the ocean background to calculate the excess brightness upon which the rain retrieval is based. The algorithm outputs two products, namely: an earth-located instantaneous rain rate by orbit revolution at 50-km resolution; and a five-day (pentad) rain rate average on a 0.5° x 0.5° latitude/longitude grid. Both products are binned in 0.5 hour universal time windows. Next, the further details of the QRad rain algorithm will be presented.

Figure 13. QRad ($T_{ex}$-R) third-order transfer function for (top) H- pol and for (bottom) V- pol. Error bars denote ± one standard deviation.
3.1. TRMM Training Data Products

The QRad rain rate algorithm was trained using a data set of sixty-six significant rain events that were observed within ±0.5 hrs with TMI. Figure 12 presents the locations of these sixty-six rain events that occurred over a nine month period in 1999 and 2000. In this algorithm development activity, we use the following TRMM products available through the TRMM Science Data and Information System (TSDIS) (http://tsdis.gsfc.nasa.gov): (1) 2A12 product, TMI derived surface rain rate over oceans, and (2) 3A11 product, TMI derived monthly freezing level over oceans.

We use the TRMM 2A12 product to provide surface rainfall rate to train the QRad rain algorithm. The 2A12 algorithm retrieves precipitation based upon nine channels of TMI brightness temperature [Kummerow et al., 1996]. This algorithm uses a Bayesian approach that utilizes cloud resolving models to generate a large database of potential hydrometeor profiles and a microwave radiative transfer model to compute the corresponding TMI channel brightness temperatures. This algorithm generates vertical hydrometeor profiles on a pixel basis. For each pixel, cloud liquid water, precipitation water, cloud ice water, precipitation ice, and the latent heating are given at 14 vertical layers. The surface rainfall and associated confidence are also computed.

We use the TRMM 3A11 product to estimate the height of the rain over the ocean for use in the QRad algorithm. The TMI 3A11 algorithm [Wilheit et al., 1991] also uses the TMI brightness temperatures to infer the freezing level, which is the estimated height of 0°C isotherm over oceans in 5° x 5° boxes for one month. It also produces 5° x 5° monthly oceanic rainfall maps using TMI Level-1 brightness temperatures.

3.2. Excess Brightness Temperature

The oceanic microwave brightness temperature when viewed through a raining atmosphere is greater than that when viewed through a clear atmosphere. Rain can be inferred from the differential (excess) part between the raining and clear ocean $T_b$; so the extraction of the rain signal depends directly upon the knowledge of the ocean

![Figure 14](image_url). Instantaneous integrated rain rate comparisons for sixty-six collocated rain events for QRad and TMI. Spatial resolution is 0.5° (50 km).
brightness when viewed through an intervening atmosphere without rain. The brightness temperature observed by the satellite microwave radiometer is determined by the electromagnetic frequency, polarization, incidence angle and by a number of atmospheric geophysical variable profiles including temperature, oxygen density, water density (vapor, cloud liquid and rain) as well as the ocean surface geophysical variables: sea surface temperature, salinity and surface wind speed. The usual remote sensing scenario is for the observing microwave radiometer to have the number of independent measurements greater than the number of unknown geophysical parameters. For example, according to Wentz and Spencer [1998], they use 7 SSM/I channels to retrieve 4 parameters; surface wind speed, integrated water vapor, integrated cloud liquid water and path average rain rate. Parameters that contribute significantly to the brightness but are not retrieved are known a priori, frequently from climatology or numerical models.

[26] Mears et al. [2000] have characterized the monthly mean ocean Tb for the QRad channels using seven years of measurements from the SSM/I. This ocean brightness temperature climatology accounts for all of the geophysical parameters except the transient effects of rain and surface winds (which have been removed in the data analysis). Fortunately, the Ku-band (13.4 GHz) Tb responds weakly to the atmospheric and surface geophysical parameters included in this climatological background. Further, all of these parameters vary slowly in space and time (seasonally). As an example, the dynamic range of the horizontally polarized ocean background temperature with latitude for the month of March is (91 K ~ 103 K), while the vertically polarized ocean background temperature for the same month lies in the range (165 K ~ 182 K). In both cases, the longitude variations are almost flat.

Table 6. Instantaneous Integrated Rain Rate Differences for Six TMI Ranges

<table>
<thead>
<tr>
<th>TMI Range</th>
<th>Number of Points</th>
<th>Difference Mean</th>
<th>Difference Std.</th>
<th>Difference rms/(TMI)</th>
</tr>
</thead>
<tbody>
<tr>
<td>0–4</td>
<td>2498</td>
<td>1.832</td>
<td>3.420</td>
<td>2.519</td>
</tr>
<tr>
<td>4–8</td>
<td>984</td>
<td>0.410</td>
<td>4.896</td>
<td>0.842</td>
</tr>
<tr>
<td>8–12</td>
<td>683</td>
<td>-0.340</td>
<td>6.469</td>
<td>0.659</td>
</tr>
<tr>
<td>12–24</td>
<td>922</td>
<td>-1.983</td>
<td>10.642</td>
<td>0.636</td>
</tr>
<tr>
<td>24–32</td>
<td>302</td>
<td>-2.660</td>
<td>15.654</td>
<td>0.573</td>
</tr>
<tr>
<td>&gt;32</td>
<td>476</td>
<td>-1.356</td>
<td>37.299</td>
<td>0.700</td>
</tr>
</tbody>
</table>

aUnits are in km*mm/hr. For each range, the difference (QRad minus TMI) is calculated.

Figure 15. Probability density function for integrated rain rate at sixty-six collocated rain events for QRad and TMI.
On the other hand, rain and wind are very transient with weather systems, and they need to be retrieved simultaneously. Because there are only two QRad channels (V- and H-pol), we use the ocean (and atmosphere) brightness climatology as a priori information with collocated wind speed measurements provided by the SeaWinds [QuikSCAT, 2001]. We define the “excess brightness” \( T_{ex} \) as the residual of the average measured QRad \( T_b \) after subtracting ocean background brightness temperature (which includes non-raining atmosphere) and the brightness temperature due to the surface wind speed. Thus the polarized \( T_{ex} \) is

\[
T_{ex} = T_{QRad} - T_{ocean} - T_{bw.speed} \tag{3}
\]

where

\[
T_{QRad} = \frac{1}{n} \sum_{i=1}^{n} T_{bi}, \quad \text{is the average measured QRad } T_b, \quad \text{K} \tag{4}
\]

where \( n \) is the number of pulses within a gridded measurement, \( T_{ocean} \) is the ocean background \( T_b \), K (includes atmosphere without rain), \( T_{bw.speed} \) is the \( T_b \) due to the wind speed, K, and \( p \) is the polarization.

The ocean background is interpolated to the day of the observation using monthly latitude/longitude tables at 0.5° spatial resolution. The QSCAT L2B ocean surface wind speeds (JPL Physical Oceanography Distributed Active Archive Center (PO.DAAC) web site: http://podaac.jpl.nasa.gov/quickscat/) are derived from the same receiver measurements, as the QRad \( T_b \) measurements, and therefore they are perfectly collocated. The theoretical brightness temperature due to wind speed [Wang, 2001] is calculated as:

\[
(T_{bw.speed})_p = a_0 + a_1 * wspd + a_2 * wspd^2 + a_3 * wspd^3, \quad \text{K} \tag{5}
\]

where \( a_{ip} \) are empirical wind speed coefficients (\( p = V\)-pol and H-pol) given in Table 4 and \( wspd \) is the QSCAT collocated ocean surface wind speed, m/s.

Typical values for the polarized brightness temperature \( (T_{bw.speed})_p \) due to a wind speed measurement of 8 m/s are 2.5 K for the H-pol and 1.2 K for the V-pol. However, for ocean winds between 0–10 m/s and in the presence of rain, QSCAT wind retrievals are typically 10–15 m/s independent of the true wind speed. Thus, when rain is present, an erroneous wind speed correction is made, which biases the excess brightness temperature low (\( \sim 5–10 \)K). Fortunately, this \( T_b \) error is compensated during the development of the empirical \( T_b \)-R relationship.
3.3. Excess Brightness - Integrated Rain Rate Relationship

The rain rate algorithm is a statistical based retrieval that uses an empirical brightness temperature - rain rate (Tb-R) relationship. This relationship is derived using a QRad brightness temperature and TMI integrated rain rate data set from sixty-six significant rain events that are observed within ±0.5 hrs. In the propagation direction, the total atmospheric absorption and emission of microwave energy is directly proportional to the rain path length; thus the observed rain brightness temperature is proportional to the integrated rain rate.

The Tb-R relationship is calculated using a regression analysis of the QRad excess brightness (T ex) with the corresponding collocated TMI integrated rain rate (IRR). First QRad T ex are produced on 0.25° grid, and the corresponding TMI 2A12 surface rain rates are converted to IRR. Because the TMI integrated rain rate value is not available in 2A12, the IRR is approximated to be the product of the TMI surface rain rate (mm/hr) and the rain path length (km). For this calculation, we use the TMI retrieved freezing level (TMI 3A11 product) as the rain height interpolated to 0.5° spatial resolution and multiply by the secant of the TMI incident angle (52.8°). For example, a typical average value for rain height near the equator during the month of March 2000 is about 4.9 km.

Next, the T ex and the IRR are averaged over a 0.5° × 0.5° earth grid that corresponds to the effective resolution of the QRad antenna. In this manner, we transfer the beam-fill correction from TMI to QRad. At high rain rates associated with small convective rain cells, the beam-fill correction does not scale well, thus QRad rain rates will be significantly underestimated. Finally, these data are binned by TMI IRR, averaged and then used in a least-squares curve fit procedure to determine an optimal 3rd order polynomial. This polynomial is forced to pass through the origin, which produces a T ex-R function with odd symmetry about zero T ex. This odd function regression is adopted to cancel (in the mean) the effect of the QRad measurement noise (ΔT) that frequently causes the T ex to be negative at low rain rates. The estimated coefficients along with their respective estimated standard errors, are provided in Table 5, and Figure 13 shows the resulting transfer function with error bars of ± one standard deviation for each bin.
Figure 18. Retrieved rain rate histograms from Monte Carlo simulation. Three curves are for noise-free Tb measurement (asterisks), $\Delta T = 1$ K (open circles), and $\Delta T = 5$ K (pluses).

Figure 19. QRad estimated rain rate histogram and fitted convolution probability density.
3.4. Integrated Rain Rate

The integrated rain rate is calculated from the polarized Tex using the T_b-R relationship given as:

\[
IRR_p = b_0 + b_1 T_{ex} + b_2 T_{ex}^2 + b_3 T_{ex}^3
\]

where \(b_i\) is the regression coefficients, given in Table 5.

The final rain rate is the weighted-average of the polarized rain rates. The usual procedure is to weight measurements by their inverse variances; but for QRad, the variances for V- and H-pol are similar. However, the dynamic range of the Tex’s differ by approximately a factor of two (H-pol \(\sim 2 \times\) V-pol range); therefore we weight these two rain retrievals by their dynamic ranges. Since H-pol is less affected by the QRad \(\Delta T\) noise, it is given greater weight in the final result:

\[
IRR = c_0 + c_1 \frac{2 \cdot IRR_h + IRR_v}{3} \text{ km} \cdot \text{mm/hr}
\]

where \(c_0\) is the empirically derived bias for no rain areas and \(c_1\) is the empirically derived slope that matches the TMI training data set rain accumulation. In the current version of the algorithm, \(c_0\) and \(c_1\) have values of approximately zero and unity, respectively.

However for larger IRR’s the retrievals are well behaved in the mean. This may be better examined in the statistics of differences (QRad minus TMI) presented in Table 6. For this comparison, we use the same data as Figure 14; but now we bin the data in six ranges of TMI IRR. The mean of the individual histograms is near zero, that verifies the T_b-R least mean squares regression procedure; however the standard deviations are large as a result of the poor QRad \(\Delta T\).

Additional quantitative comparisons between TMI and QRad for the sixty-six rain events are presented in terms of the IRR probability density functions (pdf’s) and cumulative distribution functions (cdf’s) shown in Figures 15 and 16, respectively. Note that in Figure 15, only rain rates greater than zero are presented for TMI. The large \(\Delta T\) causes some distortion in the QRad pdf especially for low IRR; however, this does not produce a significant accumulation error as seen by examining the QRad cdf. This is the result of using the \(c_1\) coefficient in equation (7) to adjust QRad to match the TMI IRR accumulation.

To evaluate the hypothesis that averaging negative and positive rain rates results in the proper mean value, we examined the histograms of QRad IRR’s for 50-km pixels over large non-raining areas. For this analysis, about sixteen orbits were examined and many large regions at least 10° \(\times\) 10° were selected where there was apparently no rain, which resulted in about 10,000 pixels. Histograms were examined individually and collectively with similar results as presented in Figure 17. The mean IRR of these combined non-raining areas is nearly zero (\(-0.355 \text{ km} \cdot \text{mm/hr}\)). We use this small offset in the QRad rain retrieval algorithm,
bias $c_n$ in equation (7), to make the average instantaneous rain product zero mean for non-raining areas.

[38] Since the QRad rain rate algorithm is applicable only over the ocean, we use a conservative land mask with extended land boundaries (and small islands deleted) to determine where the rain rate algorithm is applied. Unfortunately, when QRad measurements are close to land, the measured $T_b$ is also affected by the “hot” radiance from land that enters through the antenna pattern side-lobes. Thus, within about 150 km of land, the measured QRad $T_b$ has a land bias of about +5 to +10 K. To compensate for this effect, the background brightness temperature over the land is set to its typical value of 270K, and the ocean/land background is smoothed using a 3 x 3 pixel window to eliminate the effect of the sharp land-ocean boundary. In this way, the ocean brightness temperature near the boundary is elevated in an attempt to remove the influence of land on $T_{ex}$. The final step is to evaluate the monthly rain rate at all land/water boundaries and identify anomalous negative rain rates, when ocean background is too high; and positive rain rates, when ocean background is too low. The final land mask is subjectively adjusted to remove these anomalous rain rates that may result along the land borders.

3.5. Rain Retrieval Errors

[39] There are several sources of error in the retrieved rain rates; but the one that predominates is the random component of the QRad brightness temperature measurement error. Because of this larger than normal $\Delta T$, the excess brightness temperature includes a large random, zero-mean, Gaussian noise component that distorts the retrieved rain rate pdf and even produces unrealistic negative rain rates. We believe that after spatial and temporal averaging of both positive and negative rain rates, the majority of this noise will

Figure 21. Example of instantaneous rain rate images produced by QRad and TMI. Spatial sampling is 0.125° (12.5 km), and coincidence time difference is ~20 min. See color version of this figure at back of this issue.
cancel, and the result will be a reasonable estimate of the true average rain rate. For non-raining regions, the average of the negative and positive rain rates will be zero (after a small bias is removed); and for raining regions, the average approaches the true rain rate. This is the fundamental premise of our rain rate retrieval. To assess the effect of random measurement $\Delta T$ on the rain rate retrieval, convolution models are applied both in the forward and inverse directions.

First, a Monte Carlo simulation was performed in the forward direction [Wang, 2001]. The TMI measured rain rate density was assumed to be “nature”, and the “noise-free” excess brightness temperatures were generated using the inverse $T_b$-R relationship. Next, noisy $T_{ex}$ were created by adding Gaussian noise, and then converted to rain rate using the $T_b$-R relationship. Results for noise-free and noisy rain retrievals are presented in Figure 18 for $\Delta T$’s of 1 K and 5 K. The $\Delta T = 5$ K simulation is representative of the QRad $T_{ex}$’s after averaging over the $0.5^\circ \times 0.5^\circ$ grid. As expected, the noisy rain retrieval density functions are the convolution of the Gaussian measurement error pdf with the noise-free rain rate density. The effect of the QRad measurement $\Delta T$ is most obvious at low rain rates where the approximately exponential noise-free density function is significantly reduced in amplitude and broadened. Further, physically unrealistic negative rain rates are produced.

Next, a convolution model is applied to the estimated QRad rain rates to deconvolve the measurement noise and rain rate populations. The analytical form of the convolution probability density function can be readily found from the analytical forms of the Gaussian and exponential components via the convolution integral formula. Once the overlaid convolution pdf has obtained a good fit to the empirical data, the component distributions can be resolved and

![Figure 22. Instantaneous rain rate comparisons for a hundred and eight collocated rain events for QRad and TRMM 3B42RT HQ (TMI and SSM/I) product. Spatial resolution is 0.25° (25 km) and coincidence time difference is <75 min.](image)

<table>
<thead>
<tr>
<th>HQ Range</th>
<th>Number of Points</th>
<th>Difference Mean</th>
<th>Difference Std.</th>
<th>Difference rms/(HQ)</th>
</tr>
</thead>
<tbody>
<tr>
<td>0–1</td>
<td>14582</td>
<td>0.5330</td>
<td>0.9548</td>
<td>2.6428</td>
</tr>
<tr>
<td>1–2</td>
<td>7842</td>
<td>0.3571</td>
<td>1.7383</td>
<td>1.2278</td>
</tr>
<tr>
<td>2–4</td>
<td>6492</td>
<td>0.1350</td>
<td>2.7590</td>
<td>0.9736</td>
</tr>
<tr>
<td>4–8</td>
<td>3637</td>
<td>−0.5062</td>
<td>4.0268</td>
<td>0.7385</td>
</tr>
<tr>
<td>&gt;8</td>
<td>1939</td>
<td>−4.2267</td>
<td>8.9058</td>
<td>0.7218</td>
</tr>
</tbody>
</table>

*Units are in mm/hr. For each range, the difference (QRad minus HQ) is calculated.
examined. The fit is obtained simply by varying the unknown parameters of the convolution density. Note that the convolution density is parameterized by the constituent pdf parameters; specifically, the mean and variance of the Gaussian distribution and the mean (or shape parameter) of the exponential distribution. Figure 19 shows the histogram of QRad estimated rain rates developed from equation (7) as well as a candidate fitted convolution density of a Gaussian and exponential. Although the convolution model provides a good fit for small rain rates, it underestimates the proportion of data in the right tail, i.e. rain rates larger than 15 km*mm/hr or more. Figure 20 shows the subsequent constituent distributions compared to the TMI rain rates which are taken to be “nature”. Clearly, the exponential model is inadequate in that it does not roll off fast enough to fully capture the tail behavior of the rain rate distribution. However, the “smearing” effect and relative size of the QRad measurement $D_T$ can be obtained by examination of the estimated Gaussian distribution; specifically the estimated standard deviation which was found to be 3.5 km*mm/hr in this case. This estimate provides a quantitative way to assess the effect of the measurement $D_T$ on the QRad rain rate estimation procedure.

[42] It is readily observed that the convolution models in both the forward and inverse directions fail to exhibit an adequate fit in the tail of the empirical distribution of estimated (QRad derived) and calibration (TMI) rain rates. However, these examples illustrate that once a sufficient pdf model can be developed to model the rain rate distribution, the convolution model does have merit in illustrating the effect of measurement $\Delta T$ on the QRad estimates. Further investigation into more appropriate rain rate pdf models as well as objective parameter estimation techniques are currently being developed. Other secondary sources of error are the result of the following:

[43] 1. Convective and stratiform rain type differences. For the same rain rate, different rain types can produce differences in brightness temperatures of order a few 10’s K, which is neglected in the $T_b$-R relationship. To compensate for this effect, the QRad/TMI training data set was selected over a range of geographic locations and seasons to produce an average $T_b$-R relationship based upon the convective/stratiform conditions encountered.

[44] 2. Beam-fill differences between QRad and TMI due to antenna spatial resolution. The SeaWinds scatterometer antenna was designed to optimize the antenna boresight gain at the expense of the antenna main beam efficiency. As a result, the QRad antenna collects energy over an effective surface area that is approximately twice that of the TMI (50 km sampling for QRad compared to 25 km sampling for TMI). To produce the QRad $T_b$-R relationship, the TMI rain rates were averaged over 50 km to match the QRad $T_b$.
measurement resolution; but the empirical beam filling corrections applied to TMI do not scale linearly. This will result in a systematic underestimation of QRad peak rain rates compared to the TMI measurements.

3. Long term systematic radiometric calibration drift. The QRad calibration stability illustrated in Figure 10 shows an rms variation of 1.4 K. This effect contributes to uncertainty in the empirically derived coefficients used in the retrieval algorithm and to biases in the average retrieved rain rates. Nevertheless, this error source is considered secondary to the error introduced by the large QRad $\Delta T$.

4. QSCAT wind speed retrieval errors and the resulting reduction in excess brightness temperatures. In the

Figure 24. Examples of rain events measured by QRad (right) and TRMM 3B42RT HQ (TMI and SSM/I) product (left). Spatial resolution is 0.25° (25 km) and coincidence time difference <35 min. See color version of this figure at back of this issue.
presence of rain (and at low to moderate ocean wind speeds), QSCAT wind retrievals are bogus, typically 10–15 m/s independent of the true wind speed. This error can lower the $T_{\text{ex}}$ by 5–10 K (worst case); however this effect is largely compensated by the empirical Tb-R relationship regression; thus this is not considered to be a significant source of rain retrieval error.

Figure 25. Other examples of rain events measured by QRad (right) and TRMM 3B42RT HQ (TMI and SSM/I) product (left). Spatial resolution is 0.25° (25 km) and coincidence time difference <60 min. See color version of this figure at back of this issue.
images. This is supported by good comparisons between the ocean background and the three-day brightness temperatures used in the QRad external $T_b$ calibration. This is significant because the QRad effective brightness derived from TMI still has transient effects of winds (rain flags remove transient effects of rain) in addition to inter-annual variability of other variables from climatology. Moreover, in the future, we plan to investigate the utility of improving the

**Figure 26.** Monthly rain images produced by QRad, TMI and SSM/I F13 for March 2000. Spatial resolution $0.5^\circ$ (50 km).

**Figure 27.** Monthly, global, $0.5^\circ 	imes 0.5^\circ$ spatially averaged, rain rate differences for March 2000. From the left are QRad-TMI, SSMI-TMI, and QRad-SSMI.
algorithm by using a microwave radiative transfer model to calculate the daily ocean background instead of using the $T_b$ climatology. For this approach, we would use the daily averaged satellite measurements of SST, water vapor and cloud liquid water to remove the inter-annual climatology variability.

Error in estimating the integrated rain rate. Because the TMI integrated rain rate value is not available, the IRR is approximated to be the product of the TMI surface rain rate (mm/hr) and the rain path length (km). Since both the surface rain rate from the TMI 2A12 product and the rain height from the TMI 3A11 product have random errors, this produces increased error in the $T_b$-R empirical relationship. However, the radiometer excess brightness temperature depends upon the integrated rain rate along the propagation path; and because the height of rain varies significantly over latitude, we believe that using this IRR is the best compromise. Further, the TMI training set is distributed over the full latitude range of TRMM, which provides an averaging effect. However, since the QRad algorithm is also applied beyond the latitudinal range of TRMM, caution is advised because of the unknown accuracy in these regions.

4. Validation of QRad Rain Retrievals

4.1. Validation Data Products

TMI, an improved design of the SSM/I instrument, is dedicated to obtaining quantitative measurements of rainfall. The oceanic instantaneous rain rate, measured by the TMI is widely accepted by the science community to be the best estimate of the true rain rate available from a passive microwave sensor. Thus, to evaluate the QRad retrieved rain rate capabilities, we use the TRMM 2A12 instantaneous surface rain rate and the TRMM 3B42RT surface rain rates for the comparison data set. The TMI 2A12 instantaneous rain rate product has been validated by the TRMM science team through numerous comparisons with other independent rain measurements [Kummerow et al., 2000].

Table 8. Monthly Average Rain Rate Differences Between QRad/TMI, SSMI/TMI, and QRad/SSMI for March 2000*

<table>
<thead>
<tr>
<th>Difference</th>
<th>Number of Points</th>
<th>Difference Mean</th>
<th>Difference Std.</th>
</tr>
</thead>
<tbody>
<tr>
<td>QRad – SSMI</td>
<td>75463</td>
<td>9.877 e-2</td>
<td>2.50 e-1</td>
</tr>
<tr>
<td>QRad – TMI</td>
<td>75463</td>
<td>1.148 e-1</td>
<td>2.78 e-1</td>
</tr>
<tr>
<td>SSMI – TMI</td>
<td>75463</td>
<td>1.892 e-2</td>
<td>2.56 e-1</td>
</tr>
</tbody>
</table>

*Units are in mm/hr.

4.2. Instantaneous Rain Rates

A typical instantaneous rain image example is given in Figure 21. The upper panel shows the TMI/QRad near-simultaneous overlapping swaths. Both satellites were in descending revs and observed the rain event with a 20 minute pass time difference. The corresponding rain images are given in the lower panel. For clarity of presentation, both rain images were resampled to 0.125° resolution. The color bar on the right side indicates the rain rate values (mm/hr); and both rain images have identical color scales for retrieved rain rates. The shape and intensity of the rain event were well captured by QRad. In fact, the correlation coefficient for the two images is found to be 89.7%.

Additional evaluations of the instantaneous QRad retrieval algorithm consisted of comparisons with the high quality merged TRMM 3B42RT real time multi-satellite precipitation data product. A hundred and eight significant rain events that were observed by QRad and HQ microwave radiometers are used as an additional independent data set for this validation activity. Overall the rain intensity and spatial rain patterns were well captured by QRad and the correlation coefficients between corresponding rain images was typically >0.70.

The first quantitative comparison for these hundred and eight rain events is presented as a scatter plot in Figure 22. Statistical results of the differences (QRad minus HQ) are presented in Table 7, where we bin the data in five ranges of HQ rain rate. Although the standard deviations for the individual bins are large due to the poor QRad $\Delta T$, the retrievals are well behaved in the mean.
The second quantitative comparison for the hundred and eight rain events is presented in terms of the rain rate probability density functions (pdf’s) shown in Figure 23. Clearly, the large ΔT causes some distortion in QRad pdf for low rain rate values; however, for larger rain rates >2.5 mm/hr, the QRad pdf captures the behavior of the HQ rain rate distribution.

Next, we present sample image comparisons of collocated rain events of QRad and HQ retrieved rain rates. Although these collocated rain events are obtained from the 3-hour UTC windows, we utilize a satellite orbit database, along with specialized collocation tools to estimate the overpass time differences between QRad and HQ observations. These collocations span a period of about two weeks during the month of June 2003. First, the QRad rain was put into 3-hour universal time windows (±90 minute span around synoptic observation hours 00 UTC, 03 UTC, 06 UTC, ..., 21 UTC). Then, the resulting time binned rain images were smoothed by resampling to a 0.25° latitude/longitude grid to match the HQ rain product resolution.

The upper panel in Figure 24 shows a collocated rain event with low rain values that was observed on June 18, 2003 during the 06 UTC time-window where the coincidence time differences are <35 min. The QRad rain rates are shown on the right side, while the HQ rain rates are shown on the left side, and the color bars indicate the rain rate (mm/hr) values. The correlation coefficient between the two images is 85%. The lower panel shows a second collocated rain event with moderate rain values that was observed on June 21, 2003 during the 06 UTC time-window where the coincidence time differences are also <35 min. For this case, the spatial correlation coefficient is 75%. A third rain image comparison presented in the upper panel of Figure 25 represents an example of high rain rate that was observed on June 24, 2003 during the 15 UTC time-window where the coincidence time differences are <60 min. The correlation is found to be 80%. The last collocated rain event example is shown in the lower panel of the same figure. This rain event was observed on June 25, 2003 during the 15 UTC time-window where the coincidence time differences are also <60 min. The correlation coefficient for this event is found to be 85%.

In general, there is very good spatial correlation between QRad and HQ rain patterns. Because of the smaller IFOV and lower ΔT, the HQ images are “crisper”; nevertheless, the shape and relative intensity of the rain events are well captured by the QRad images. On an absolute basis, the QRad underestimates the higher rain rates because of the non-linear effects of beam filling. Further, the effects of the high ΔT result in “noisy pixels” that is apparent in the QRad rain images. Most differences between HQ and QRad are attributed to errors in the QRad retrievals; however some differences may be “real” in that they could be the result of the different pass times of QRad and HQ over the rain events.

4.3. Averaged Rain Rates

For the average rain rate product, we perform temporal (pentad) and spatial (0.5° × 0.5°) averaging of all instantaneous rain rate values (positive and negative) which significantly reduces the random component of the rain retrieval. As an example, Figure 26 shows the average rain rate for March 2000, produced from QRad, TMI and SSM/
I-F13, averaged over the global region ±40° latitude on a 0.5° × 0.5° latitude/longitude grid. As the spatial resolution decreases (i.e., spatial averaging area increases), the correlation improves. An example of the differences between the three rain rate retrievals for 0.5° × 0.5° for March 2000 is presented in Figure 27, and the statistical measures for these cases are given in Table 8. Here there is excellent agreement between TMI and SSMI and quite reasonable comparisons for both with QRad. Most of the difference occurs in the vicinity of the ITCZ area where the convective rain activity predominates.

[59] In general, there is excellent correlation between the spatial patterns of rain; however there are fine scale differences due to the larger spatial resolution of QRad, and its

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**Figure 29.** Typical examples of near-simultaneous collocation cases for QRad (right) and TRMM 3B42RT VAR (visible and infrared) product (left). Spatial resolution is 0.25° (25 km). See color version of this figure at back of this issue.
poorer radiometric precision (AT). Nevertheless, the shape and the relative intensity of the rain are well captured by QRad.

Finally, Figure 28 shows a time series of QRad and TMI zonal five-day (pentad) rain rates, averaged over the tropical ocean from 0°N to 20°N. Pentad averages were calculated for about nine months during January 2000 through September 2000. Although QRad slightly overestimates the rain rate, there is high correlation between these two time series (≈86%), and this result is in excellent agreement with a similar study of Imaoka and Spencer [2000] between pentad averages for TMI and SSMI.

5. Summary

This paper discussed the details of the QRad statistical rain retrieval algorithm. Comparisons between rain products derived from QRad, and rain retrievals obtained from independent microwave rain measuring instruments are presented. Results demonstrate that QRad rain measurements are in very good agreement with these independent rain estimates in terms of the spatial distribution of the rain patterns and the relative rain intensity. However, due to the poor radiometric resolution (ΔT) of QRad, some fine scale differences between the rain retrievals are noticed.

When compared to rain measurements obtained from visible and infrared satellite observations, QRad rain estimates perform superbly. As an example, Figure 29 presents two collocated rain events between QRad and the TRMM 3B42RT VAR data product. The QRad rain rates are shown on the right side, while the VAR rain rates are shown on the left side. The color bars are proportional to the rain rate (mm/hr) values. For these comparisons we apply a threshold of 1mm/hr to QRad rain rates to eliminate any random bogus rain pixels. In both cases, it can be seen that the VAR rain estimates failed to detect a significant portion of the low and moderate rain event structure. These examples are quite typical, and they emphasize the superior performance of the microwave rain retrievals compared to rain estimates from visible and infrared sources.

The major scientific utility of QRad rain measurements is that they provide additional independent temporal and spatial sampling of the oceanic rain, which complements the coverage provided by TMI and the SSMIs’ instruments. Thus the QRad rain time series from 1999 to present is a valuable addition to the ocean precipitation climatology data set that can be potentially used to improve the diurnal estimation of the global rainfall, which is a goal for the future Global Precipitation Mission program. Moreover, the early availability of QRad data will afford users early access to learn to use less-precise rain measurements that will occur in the future with the use of less-capable constellation satellites. Finally, these QRad rain estimates will be available in the planned data reprocessing (FY 2006) to users of QuikSCAT winds to improve the rain flagging of rain-contaminated oceanic wind vector retrievals.

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References


Figure 2. Ocean sampling, daily average revisit time. (top) TMI and 3-SSMI’s and (bottom) sampling with QRad added.

Figure 21. Example of instantaneous rain rate images produced by QRad and TMI. Spatial sampling is 0.125° (12.5 km), and coincidence time difference is ~20 min.
Figure 24. Examples of rain events measured by QRad (right) and TRMM 3B42RT HQ (TMI and SSM/I) product (left). Spatial resolution is 0.25° (25 km) and coincidence time difference <35 min.
Figure 25. Other examples of rain events measured by QRad (right) and TRMM 3B42RT HQ (TMI and SSM/I) product (left). Spatial resolution is 0.25° (25 km) and coincidence time difference <60 min.
Figure 29. Typical examples of near-simultaneous collocation cases for QRad (right) and TRMM 3B42RT VAR (visible and infrared) product (left). Spatial resolution is 0.25° (25 km).